

Alexandrov's theorem in curved manifolds

Simon Brendle

May 11, 2012

The isoperimetric problem

Queen Dido's problem: Minimize the length among all embedded curves in the plane with a given enclosed area.

Generalize this to higher dimensions: Minimize the area of an embedded hypersurface in \mathbb{R}^n among all surfaces that enclose a given volume V .

Theorem. *Round spheres are optimal for the isoperimetric problem. More precisely, if Ω is a domain in \mathbb{R}^n satisfying $\text{vol}(\Omega) = \text{vol}(B)$, then $|\partial\Omega| \geq |\partial B|$.*

Symmetrization techniques show that if a minimizer exists it must be a sphere. Existence of a minimizer is a non-trivial issue. A rigorous proof (based on Brunn-Minkowski inequality) was found in the 20th century.

The first variation of surface area

Study the isoperimetric problem using the Calculus of Variations developed by Bernoulli, Euler, and others.

Consider a domain $\Omega \subset \mathbb{R}^n$ with smooth boundary. Let X be a smooth vector field, and let $\Omega_t = \varphi_t(\Omega)$, where φ_t is the flow generated by X . Then

$$\left. \frac{d}{dt} \text{vol}(\Omega_t) \right|_{t=0} = \int_{\partial\Omega} \langle X, \nu \rangle$$

and

$$\left. \frac{d}{dt} |\partial\Omega_t| \right|_{t=0} = \int_{\partial\Omega} H \langle X, \nu \rangle.$$

Here, H is the mean curvature of the surface $\partial\Omega$ (can be viewed as an L^2 gradient of the area functional). Note that $H = \lambda_1 + \dots + \lambda_{n-1}$ where the λ_i 's are the principal curvatures of $\partial\Omega$ (i.e. the reciprocals of the curvature radii).

Surfaces with $H = 0$ are minimal surfaces (soap films). The catenoid in \mathbb{R}^3 is the simplest example.

Surfaces with constant mean curvature in Euclidean space

Proposition. *Suppose that the first variation of area is zero for variations that leave the enclosed volume unchanged. Then the mean curvature of $\partial\Omega$ is constant: $H = c$.*

The constant c can be viewed as a Lagrange multiplier.

Problem: Analyze hypersurfaces in \mathbb{R}^n with constant mean curvature.

Theorem (Alexandrov 1956). *Let Σ be a hypersurface in \mathbb{R}^n which has constant mean curvature and is embedded (i.e. no self-intersections). Then Σ is a round sphere.*

Wente, Kapouleas: There exist non-trivial examples of constant mean curvature surfaces, but these fail to be embedded.

The proof of Alexandrov's theorem is based on the **method of moving planes**: Reflect Σ across a hyperplane P . Move the hyperplane P until the reflected surface touches the original surface Σ . The reflected surface must coincide with the original surface by the strict maximum principle.

Surfaces with constant mean curvature in Riemannian manifolds

Replace the ambient space \mathbb{R}^n by a curved Riemannian manifold. This question is motivated in part by questions in relativity:

- Christodoulou, Yau (1986): Let $\Sigma \approx S^2$ be a stable surface of constant mean curvature in a 3-manifold (M, g) with positive scalar curvature. Then Σ has nonnegative Hawking mass.
- Huisken, Yau (1996): Let (M, g) be an asymptotically flat 3-manifold which has positive mass (i.e. (M, g) looks Euclidean near infinity). Then (M, g) admits a foliation by constant mean curvature surfaces near infinity. Under mild extra assumptions, this foliation is unique.

- Qing-Tian, Eichmair-Metzger: Optimal uniqueness theorem for large surfaces of constant mean curvature.
- Rigger: Similar results hold when (M, g) is asymptotic to hyperbolic space near infinity. (This corresponds to a universe with negative cosmological constant.)

The Schwarzschild manifold

Recall: The Schwarzschild spacetime describes a static black hole in general relativity.

The **Schwarzschild manifold** is defined as a $\{t = \text{constant}\}$ slice in the Schwarzschild spacetime. The resulting manifold is $M = \{x \in \mathbb{R}^n : |x| \geq 1\}$, and the metric is given by

$$g = (1 + |x|^{2-n})^{\frac{4}{n-2}} (dx_1^2 + \dots + dx_n^2).$$

Note that (M, g) is asymptotic to Euclidean space near infinity. The boundary $\partial M = \{x \in \mathbb{R}^n : |x| = 1\}$ is a minimal surface and is referred to as the **horizon**.

A version of Alexandrov's Theorem for the Schwarzschild manifold

Goal: Classify all surfaces of constant mean curvature in the Schwarzschild manifold (not just the ones near infinity).

Theorem (B. 2011). *Let Σ be a closed, embedded, orientable surface in the Schwarzschild manifold with constant mean curvature. Then Σ is a sphere of symmetry, i.e. $\Sigma = \{|x| = r\}$ for some constant r .*

No assumptions on the topology of Σ or the dimension are needed. No stability assumptions are required.

Sketch of proof

Main difficulty: Method of moving planes no longer works!

New approach: Let $M = \{x \in \mathbb{R}^n : |x| \geq 1\}$ and

$$g = (1 + |x|^{2-n})^{\frac{4}{n-2}} (dx_1^2 + \dots + dx_n^2).$$

Consider the static potential

$$f = \frac{1 - |x|^{2-n}}{1 + |x|^{2-n}}$$

and the position vector field

$$X = \sum_{i=1}^n x_i \frac{\partial}{\partial x_i}.$$

Note that the $(n + 1)$ -dimensional Lorentzian metric $-f^2 dt^2 + g$ solves the **Einstein vacuum equations!** Moreover, $D_i X_j = f g_{ij}$, so X is conformal.

The main theorem is a consequence of the following proposition:

Proposition. *Let Σ be a closed, embedded, orientable hypersurface with positive mean curvature. Then*

$$(n - 1) \int_{\Sigma} f \, d\mu = \int_{\Sigma} H \langle X, \nu \rangle \, d\mu$$

and

$$(n - 1) \int_{\Sigma} \frac{f}{H} \, d\mu \geq \int_{\Sigma} \langle X, \nu \rangle \, d\mu.$$

Moreover, if equality holds then Σ is a sphere of symmetry.

The first statement follows from the divergence theorem. The second statement is related to a classical inequality of Heintze-Karcher. To prove it, we distinguish two cases:

Case 1: Σ is null-homologous.

Case 2: Σ is homologous to the horizon ∂M .

We will focus here on Case 2, which is the more difficult one.

A geometric evolution equation

We claim:

$$(n - 1) \int_{\Sigma} \frac{f}{H} d\mu \geq n \int_{\Omega} f + A(n).$$

Here, Ω is the region bounded by Σ and the horizon ∂M . $A(n)$ is a positive constant which reflects the contribution of the horizon.

Idea of proof: Deform the surface Σ with speed $-f\nu$, where ν is the outward pointing unit normal. This gives a one-parameter family of surfaces Σ_t , $t \geq 0$.

The evolution of the mean curvature is described by the equation

$$\frac{\partial H}{\partial t} = \Delta_{\Sigma} f + f \sum_i \lambda_i^2 + f \operatorname{Ric}(\nu, \nu),$$

where the λ_i 's are the principal curvatures.

Key observation: Since the Lorentzian metric $-f^2 dt^2 + g$ satisfies Einstein's equation, the potential f satisfies

$$(\Delta f)g - D^2 f + f \text{Ric} = 0.$$

Hence, the evolution of the mean curvature becomes

$$\begin{aligned} \frac{\partial H}{\partial t} &= \Delta_{\Sigma} f + f \sum_i \lambda_i^2 + f \text{Ric}(\nu, \nu) \\ &= \Delta f - (D^2 f)(\nu, \nu) + f \text{Ric}(\nu, \nu) \\ &\quad - H \langle \nabla f, \nu \rangle + f \sum_i \lambda_i^2 \\ &\geq -H \langle \nabla f, \nu \rangle + \frac{1}{n-1} f H^2. \end{aligned}$$

This allows us to show that the quantity

$$Q(t) = (n-1) \int_{\Sigma_t} \frac{f}{H} d\mu - n \int_{\Omega_t} f$$

is **monotone decreasing!**

We next analyze the limit of $Q(t)$ as $t \rightarrow \infty$: For t large, Σ_t is close to the horizon, and we obtain

$$\limsup_{t \rightarrow \infty} \int_{\Sigma_t} \frac{f}{H} d\mu \geq A(n).$$

Consequently,

$$\begin{aligned} (n-1) \int_{\Sigma} \frac{f}{H} d\mu - n \int_{\Omega} f \\ = Q(0) \geq \limsup_{t \rightarrow \infty} Q(t) \geq A(n), \end{aligned}$$

completing the proof of the Theorem.

Concluding remarks:

- The surfaces Σ_t sweep out a **lightlike** hypersurface in the Schwarzschild spacetime (i.e. the surfaces Σ_t move with the speed of light).
- The surfaces Σ_t may not be smooth (due to focal points). To overcome this problem, approximate Σ_t by smooth surfaces.

Constant mean curvature surfaces in warped product manifolds

Goal: Extend the previous analysis to more general spaces with rotational symmetry.

Consider a metric of the form $g = dr^2 + h(r)^2 g_{S^{n-1}}$. Assume that the warping function $h : [0, \bar{r}) \rightarrow \mathbb{R}$ satisfies the following structure conditions:

(H1) $h'(0) = 0$ and $h''(0) > 0$.

(H2) $h'(r) > 0$.

(H3) The scalar curvature of g is monotone decreasing in r .

Theorem (B. 2011). *Let (M, g) be a rotationally symmetric manifold satisfying conditions (H1)–(H3). Moreover, let Σ be a closed, embedded, orientable hypersurface in (M, g) with constant mean curvature. Then Σ is umbilic.*

This theorem applies, inter alia, to Euclidean space, hyperbolic space, the hemisphere, and the Schwarzschild manifold. The result also applies to the DeSitter-Schwarzschild and Anti-DeSitter-Schwarzschild manifolds (non-zero cosmological constant).

Conversely, Alexandrov's theorem fails if the scalar curvature is not monotone in r :

Theorem (F. Pacard, X. Xu). *Let g be a smooth metric on the ball $B_{\bar{r}}(0) \subset \mathbb{R}^n$ which is rotationally symmetric. Assume that the scalar curvature of g has a strict local extremum somewhere in $B_{\bar{r}}(0)$. Then there exist small spheres with constant mean curvature which are not umbilic.*

Therefore, our result is essentially optimal.

Minimal surfaces in the sphere S^3

We will now discuss a different (though related) problem:

Consider the three-dimensional unit sphere in \mathbb{R}^4 :

$$S^3 = \{(x_1, x_2, x_3, x_4) \in \mathbb{R}^4 : x_1^2 + x_2^2 + x_3^2 + x_4^2 = 1\}.$$

Note: While there are no closed minimal surfaces in \mathbb{R}^3 , the topology of S^3 allows **closed minimal surfaces**. Examples of minimal surfaces in S^3 :

- The equator $\{(x_1, x_2, x_3, x_4) \in \mathbb{R}^4 : x_1^2 + x_2^2 + x_3^2 + x_4^2 = 1, x_4 = 0\}$.
- The Clifford torus $\{(x_1, x_2, x_3, x_4) \in \mathbb{R}^4 : x_1^2 + x_2^2 = x_3^2 + x_4^2 = \frac{1}{2}\}$.

Examples of minimal surfaces in S^3 :

- H.B. Lawson, Jr. (1970): Given any positive integer g , there exists an embedded minimal surface in S^3 with genus g .
- Karcher, Pinkall, and Sterling (1988): Examples of embedded minimal surfaces in S^3 with genus 3, 5, 6, 7, 11, 19, 73, and 601.
- Further examples constructed recently by Kapouleas-Yang (doubling construction).

Moreover, Lawson has constructed an infinite family of immersed minimal tori in S^3 . These surfaces have self-intersections.

The Lawson Conjecture

Motivation: Minimal S^2 's in S^3 are unique.

Theorem (Almgren 1966). *Let $\Sigma \approx S^2$ be an immersed minimal surface in S^3 . Then Σ is congruent to the equator.*

It turns out that minimal tori in S^3 are unique as well:

Theorem (B. 2012). *The Clifford torus is the only embedded minimal torus in S^3 .*

This answers a question posed by H.B. Lawson, Jr., in 1970.

Proof of Lawson's conjecture involves an application of the maximum principle to a function which depends on a **pair of points**.

Proof of Lawson's conjecture (sketch)

Let $F : \Sigma \rightarrow S^3$ be an embedded minimal torus and let $\nu : \Sigma \rightarrow S^3$ denote the unit normal vector field. We claim that F parametrizes the Clifford torus. Since F is minimal, the sum of the principal curvatures is 0 at each point. If both principal curvatures vanish at x , we say that x is an umbilic point.

Key idea: A minimal torus in S^3 has no umbilic points. Consequently, the ratio

$$Q(x, y) = \sqrt{2} \frac{\langle \nu(x), F(y) \rangle}{|A(x)| (1 - \langle F(x), F(y) \rangle)}$$

is finite!

Consider a pair of points (\bar{x}, \bar{y}) where $Q(x, y)$ is maximal. At the point (\bar{x}, \bar{y}) , the first derivatives of Q vanish, and the second derivatives form a negative definite matrix.

Proof of Lawson's conjecture (continued)

An intricate calculation shows that $Q(\bar{x}, \bar{y}) \leq 1$. Thus, $Q(x, y) \leq 1$ for all points $x, y \in \Sigma$. An analogous argument gives $Q(x, y) \geq -1$ for all points $x, y \in \Sigma$.

On the other hand, if the gradient of the curvature is non-zero, we can arrange that $|Q(x, y)| > 1$ where x, y are very close to each other. Since $|Q(x, y)| \leq 1$ for all points $x, y \in \Sigma$, the curvature must be constant. This implies that Σ is the Clifford torus.