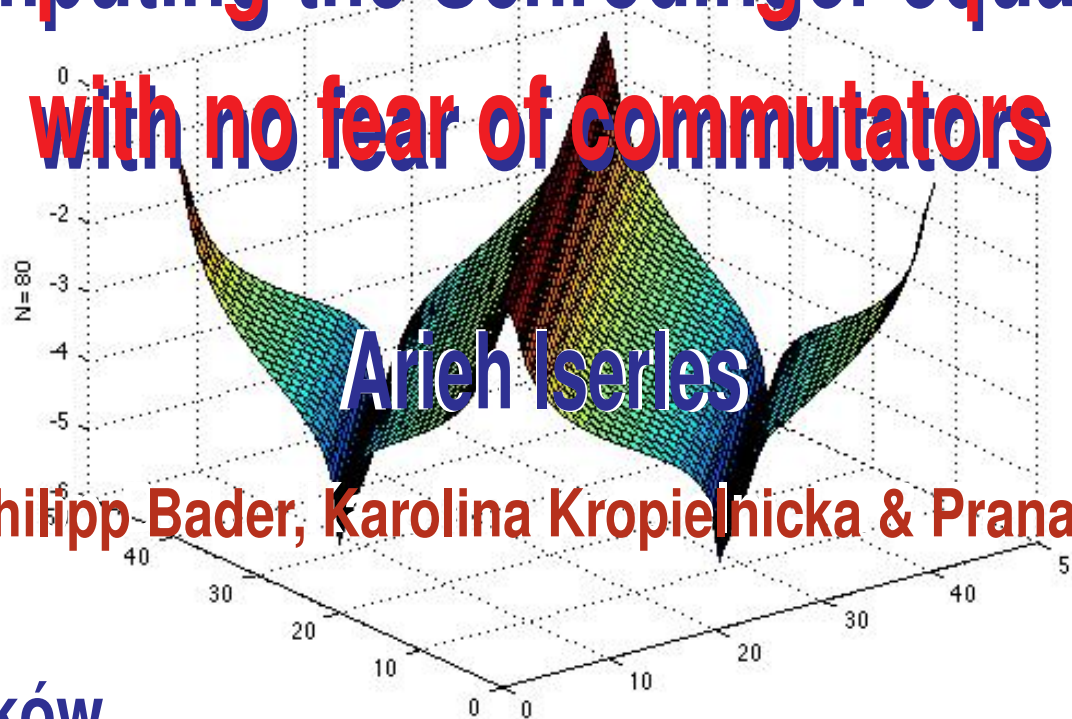


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Computing the Schrödinger equation with no fear of commutators

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6ECM, Kraków

July 2012

THE PROBLEM

Consider the linear Schrödinger equation

$$i\hbar \frac{\partial u}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 u}{\partial x^2} - V(x)u, \quad t \geq 0, \quad -\frac{1}{\sqrt{2m}} \leq x \leq \frac{1}{\sqrt{2m}},$$

where $0 < \hbar \ll 1$, and the interaction potential V is a smooth periodic function, given with smooth periodic initial and boundary conditions.

We can contemplate this equation on a multivariate torus, $i\hbar \frac{\partial u}{\partial t} = -\frac{\hbar^2}{2m} \Delta u - V(x)u$ and our theory scales up to this setting. However, our purpose here is to explore ideas, rather than presenting them in the greatest possible generality.

In physically-interesting coordinates, translating space and time,

$$\frac{\partial u}{\partial t} = i\varepsilon \frac{\partial^2 u}{\partial x^2} + i\varepsilon^{-1} V(x)u, \quad x \in [-1, 1],$$

where $0 < \varepsilon \ll 1$ but much larger than the Planck constant $\hbar \approx 1.05 \cdot 10^{-34}$: it is useful to think of the range $10^{-8} \leq \varepsilon \leq 10^{-4}$.

EXPONENTIALS AND SPLITTINGS...

The conventional approach: replace $\frac{\partial^2 u}{\partial x^2}$ by a linear combination of function values (or Fourier modes). This results in the linear system

$$\mathbf{y}' = (\varepsilon\mathcal{K} + \varepsilon^{-1}\mathcal{D})\mathbf{y}, \quad t \geq 0, \quad \mathbf{y}(0) = \mathbf{y}_0,$$

whose exact solution is

$$\mathbf{y}(t) = e^{t(\varepsilon\mathcal{K} + \varepsilon^{-1}\mathcal{D})}\mathbf{y}_0.$$

It is usual to **split** the exponential, e.g. with the **Strang splitting**

$$e^{\frac{1}{2}t\varepsilon\mathcal{K}} e^{t\varepsilon^{-1}\mathcal{D}} e^{\frac{1}{2}t\varepsilon\mathcal{K}} = e^{t(\varepsilon\mathcal{K} + \varepsilon^{-1}\mathcal{D})} + \mathcal{O}(t^3)$$

or higher-order splittings. Typically, such high-order splittings are obtained either via the **Yōšida device** or similar techniques, and result in palindromic expressions of the form

$$e^{\alpha_1 t \varepsilon \mathcal{K}} e^{\beta_1 t \varepsilon^{-1} \mathcal{D}} e^{\alpha_2 t \varepsilon \mathcal{K}} \dots e^{\beta_r t \varepsilon^{-1} \mathcal{D}} \dots e^{\alpha_2 t \varepsilon \mathcal{K}} e^{\beta_1 t \varepsilon^{-1} \mathcal{D}} e^{\alpha_1 t \varepsilon \mathcal{K}}.$$

(Palindromy implies even order and helps to preserve unitarity.)

- **Good news:** the splitting separates scales. Exponentials with ε and ε^{-1} are kept apart and the computation of each individual exponential is cheap.
- **Bad news:** we need plenty of exponentials to attain reasonable order. The number of exponentials increases *exponentially* with order and this renders high-order methods very expensive.
- **Ugly news:** the ‘scale’ of the exponentials doesn’t decrease. ideally, we would have liked to combine numerics with asymptotics, namely for the arguments of exponentials to become progressively smaller – while in the present case they are all of just two interlacing orders of magnitude.

The challenge is to develop a methodology which attains all the advantages without any disadvantages: ***A high-order splitting that separates scales and produces an asymptotic expansion in increasing powers of ε .***

ALGEBRA OF OPERATORS

The vector field is a linear combination of two linear operators: ∂_x^2 and (multiplication by) V . Let us consider the free Lie algebra \mathfrak{F} generated by ∂_x^2 and V . Note that

$$[V, \partial_x^2] = -(\partial_x^2 V) - 2(\partial_x V)\partial_x,$$

$$[[V, \partial_x^2], \partial_x^2] = (\partial_x^4 V) + 4(\partial_x^3 V)\partial_x + 4(\partial_x^2 V)\partial_x^2,$$

$$[[V, \partial_x^2], V] = -2(\partial_x V)^2,$$

$$[[[V, \partial_x^2], \partial_x^2], \partial_x^2] = -(\partial_x^6 V) - 6(\partial_x^5 V)\partial_x - 12(\partial_x^4 V)\partial_x^2 - 8(\partial_x^3 V)\partial_x^3$$

and so on. In general, $\mathfrak{F} \subset \mathfrak{G}$, where \mathfrak{G} is the Lie algebra

$$\mathfrak{G} = \left\{ \sum_{k=0}^n y_k(x) \partial_x^k : n \in \mathbb{Z}_+, y_0, y_1, \dots, y_n \text{ smooth \& periodic} \right\}.$$

$$\text{girth} \left(\sum_{k=0}^n y_k(x) \partial_x^k \right) = n$$

is the **girth** of an element in \mathfrak{G} .

PROPOSITION We have $\text{girth}([X, Y]) = (\text{girth}(X) + \text{girth}(Y) - 1)_+$ for all $X, Y \in \mathfrak{G}$.

COROLLARY Any nested commutator in \mathfrak{F} which has at least two more V s than ∂_x^2 s is necessarily zero.

For example, out of the 41 terms in the Hall basis of all the elements of \mathfrak{F} which can be written with at most 7 ‘letters’,

$$\begin{aligned} & [[V, \partial_x^2], V], V], \quad [[[[V, \partial_x^2], V], V], V], \quad [[[[[V, \partial_x^2], \partial_x^2], V], V], V], \\ & [[[[[V, \partial_x^2], V], V], V], V], \quad [[[[V, \partial_x^2], V], V], [V, \partial_x^2]], \\ & [[[[[[V, \partial_x^2], \partial_x^2], V], V], V], V], \quad [[[[[[V, \partial_x^2], V], V], V], V], V], \\ & [[[[[V, \partial_x^2], V], V], V], [V, \partial_x^2]], \quad [[[[[V, \partial_x^2], V], V], [[V, \partial_x^2], \partial_x^2]], \\ & [[[[[V, \partial_x^2], V], V], [V, \partial_x^2], V]] \end{aligned}$$

are all zero.

But this is not the real reason why all this is important! The real reason is that V never walks on its own, it is always multiplied by a large parameter ε^{-1} . The elimination of “ ε^{-1} -rich” terms allows us to derive asymptotic splittings!

SYMMETRIC ZASSENHAUS SPLITTING

We commence from the symmetric Baker–Campbell–Hausdorff formula

$$e^{tX} e^{tY} e^{tX} = e^{\text{sBCH}(t; X, Y)},$$

where

$$\begin{aligned} \text{sBCH}(t; X, Y) = & t(2X + Y) - t^3 \left(\frac{1}{24} [[X, Y], X] + \frac{1}{12} [Y, [X, Y]] \right) \\ & + t^5 \left(\frac{7}{360} [X, [X, [X, [X, Y]]] + \frac{1}{360} [Y, [Y, [Y, [X, Y]]] \right. \\ & - \frac{1}{90} [X, [Y, [Y, [X, Y]]] + \frac{1}{45} [Y, [X, [X, [X, Y]]] \\ & + \frac{1}{60} [X, [X, [Y, [X, Y]]] + \left. \frac{1}{30} [Y, [Y, [X, [X, Y]]] \right) \\ & + \mathcal{O}(t^7). \end{aligned}$$

It is possible to expand as far as we wish, noting that all the expansion terms live in the **free Lie algebra** generated by the ‘letters’ X and Y . Such algebras possess convenient bases that can be constructed algorithmically, e.g. the **Hall basis**, the **Lyndon basis** and the **Dynkin basis**.

The classical Zassenhaus splitting is

$$e^{t(X+Y)} = e^{tX} e^{tY} e^{t^2 U_2(X,Y)} e^{t^3 U_3(X,Y)} e^{t^4 U_4(X,Y)} \dots,$$

where

$$U_2(X, Y) = -\frac{1}{2}[X, Y],$$

$$U_3(X, Y) = \frac{1}{3}[Y, [X, Y]] + \frac{1}{6}[X, [X, Y]],$$

$$U_4(X, Y) = -\frac{1}{24}[[[X, Y], X], X] - \frac{1}{8}[[[X, Y], X], Y] - \frac{1}{8}[[[X, Y], Y], Y].$$

We are, however, interested in **symmetric** (i.e., palindromic) splittings. Moreover, powers of t are misleading: we need to reckon with three parameters: ε (small!), the number N of degrees of freedom once we semi-discretize (large!) and the time step Δt (small!). It makes sense to reduce this to a single currency by requiring

$$N \sim \mathcal{O}(\varepsilon^{-\rho}), \quad \Delta t \sim \mathcal{O}(\varepsilon^\sigma)$$

for some $\rho, \sigma > 0$. Note that ∂_x scales like $\mathcal{O}(N) = \mathcal{O}(\varepsilon^{-\rho})$.

In the sequel we assume that $\rho = \sigma = \frac{1}{2}$.

SYMMETRIC ZASSENHAUS

We seek a splitting of the form

$$e^{i(\Delta t)(\varepsilon \partial_x^2 + \varepsilon^{-1} V)} \approx \underbrace{e^{R_0} e^{R_1} \dots e^{R_s} e^{T_{s+1}} e^{R_s} \dots e^{R_1} e^{R_0}},$$

where (recalling in line with our ‘currency conversion’ that $\partial_x \sim \mathcal{O}(\varepsilon^{-1/2})$)

$$\begin{aligned} R_k &= R_k(\Delta t, \varepsilon, V) = \mathcal{O}(\varepsilon^{k-1/2}), \quad k = 0, 1, \dots, s, \\ T_{s+1} &= T_{s+1}(\Delta t, \varepsilon, V) = \mathcal{O}(\varepsilon^{s+1/2}). \end{aligned}$$

Let $\tau = i\Delta t = \mathcal{O}(\varepsilon^{1/2})$. Since $\tau\varepsilon^{-1}\partial_x^2 \sim \mathcal{O}(\varepsilon^{1/2})$ and $\tau\varepsilon^{-1}V \sim \mathcal{O}(\varepsilon^{-1/2})$, we commence by ‘knocking out’ the potential:

$$R_0 = \frac{1}{2}\tau\varepsilon^{-1}V, \quad Y_0 = \tau\varepsilon\partial_x^2 + \tau\varepsilon^{-1}V,$$

and rewrite the sBCH formula as

$$e^{Y_0} = e^{R_0} e^{\text{sBCH}(-R_0, Y_0)} e^{R_0}.$$

In what follows we restrict ourselves to an $\mathcal{O}(\varepsilon^{7/2})$ expansion, i.e. $s = 2$.
 After **long** calculation:

- 1 Expanding $\text{sBCH}(-\frac{1}{2}\tau\varepsilon^{-1}V, \tau\varepsilon\partial_x^2 + \tau\varepsilon^{-1}V)$ with the sBCH formula and replacing elements from \mathfrak{F} by terms from \mathfrak{G} ;
- 2 Throwing away all zero terms, but also all terms which are $\mathcal{O}(\varepsilon^\alpha)$ for $\alpha \geq -\frac{7}{2}$;
- 3 Expressing everything in the Hall basis

we obtain the **asymptotic expansion**

$$\begin{aligned} \text{sBCH}(X, Y) = & \tau\varepsilon H_1 - \frac{1}{12}\tau^3\varepsilon H_4 - \frac{1}{24}\tau^3\varepsilon^{-1}H_5 - \frac{1}{720}\tau^5\varepsilon H_{10} - \frac{1}{720}\tau^5\varepsilon^{-1}H_{11} \\ & - \frac{1}{180}\tau^5\varepsilon H_{13} - \frac{1}{288}\tau^5\varepsilon^{-1}H_{14} + \frac{1}{30240}\tau^7\varepsilon^{-1}H_{27} + \frac{1}{7560}\tau^7\varepsilon^{-1}H_{32} \\ & - \frac{101}{120960}\tau^7\varepsilon^{-1}H_{35} - \frac{89}{120960}\tau^7\varepsilon^{-1}H_{40} + \mathcal{O}(\varepsilon^{7/2}), \end{aligned}$$

where all that survives of the 41 terms in the Hall basis is

$$\begin{aligned}
H_1 &= \partial_x^2, & H_4 &= (\partial_x^4 V) + 4(\partial_x^3 V)\partial_x + 4(\partial_x^2 V)\partial_x^2, & H_5 &= -2(\partial_x V)^2, \\
H_{10} &= -\{6(\partial_x^5 V)(\partial_x V) + 12(\partial_x^4 V)(\partial_x^2 V) + 8(\partial_x^3 V)^2\} \\
&\quad - 24\{(\partial_x^4)(\partial_x V) + (\partial_x^3 V)(\partial_x^2 V)\}\partial_x - 24(\partial_x^3)(\partial_x V)\partial_x^2, \\
H_{11} &= 8(\partial_x^2 V)(\partial_x V)^2, & H_{13} &= \{2(\partial_x^5 V)(\partial_x V) - 4(\partial_x^4)(\partial_x^2 V) - 4(\partial_x^3 V)^2\} \\
&\quad + 8\{(\partial_x^4 V)(\partial_x V) - (\partial_x^3 V)(\partial_x^2 V)\}\partial_x + 4(\partial_x^3 V)(\partial_x V)\partial_x^2, \\
H_{14} &= -4(\partial_x^2 V)(\partial_x V)^2, & H_{27} &= -48(\partial_x^3 V)(\partial_x V)^3, \\
H_{32} &= 16(\partial_x^3 V)(\partial_x V)^3 + 32(\partial_x^2 V)^2(\partial_x V)^2, \\
H_{35} &= -8(\partial_x^3 V)(\partial_x V)^3 - 16(\partial_x^2 V)^2(\partial_x V)^2, & H_{40} &= -32(\partial_x^2 V)^2(\partial_x V)^2.
\end{aligned}$$

4 We next express the remaining commutators as terms in the larger Lie algebra \mathfrak{G} .

But here we have a problem!!! Ultimately, we will replace derivatives by (finite-dimensional) differentiation matrices. Even derivatives yield symmetric matrices and multiplication by i results in skew-Hermitian matrices – **good!** But, by the same token, odd derivatives \Rightarrow skew-symmetric matrices \Rightarrow (after multiplication by i) Hermitian matrices – **bad!** We must somehow get rid of odd derivatives!!!

5 Replacement of odd derivatives:

The main idea is to replace odd derivatives by linear combinations of even derivatives:

$$y(x)\partial_x = -\frac{1}{2} \int_0^x y(\xi) d\xi \partial_x^2 - \frac{1}{2} y'(x) + \frac{1}{2} \partial_x^2 \left[\int_0^x y(\xi) d\xi \cdot \right],$$

$$y(x)\partial_x^3 = -y'(x)\partial_x^2 - \frac{1}{4} \int_0^x y(\xi) d\xi \partial_x^4 + \frac{1}{4} y'''(x) - \frac{1}{2} \partial_x^2 [y'(x) \cdot]$$

$$+ \frac{1}{4} \partial_x^4 \left[\int_0^x y(\xi) d\xi \cdot \right]$$

and so on. The outcome is

$$\begin{aligned} \text{sBCH}(-R_0, Y_0) = & \tau \varepsilon \partial_x^2 - \frac{1}{12} \tau^3 \varepsilon \{ -(\partial_x^4 V) + 2\partial_x^2 [(\partial_x^2 V) \cdot] + 2(\partial_x^2 V) \partial_x^2 \} + \frac{1}{12} \tau^3 \varepsilon^{-1} (\partial_x V)^2 \\ & - \frac{1}{720} \tau^5 \varepsilon \{ 6(\partial_x^5 V)(\partial_x V) + 12(\partial_x^4 V)(\partial_x^2 V) + 4(\partial_x^3 V)^2 \\ & - 12\partial_x^2 [(\partial_x^3 V)(\partial_x V) \cdot] - 12(\partial_x^3 V)(\partial_x V) \partial_x^2 \} - \frac{1}{90} \tau^5 \varepsilon^{-1} (\partial_x^2 V)(\partial_x V)^2 \\ & - \frac{1}{90} \tau^5 \varepsilon \{ 3(\partial_x^5 V)(\partial_x V) - 2(\partial_x^4 V)(\partial_x^2 V) - 4(\partial_x^3 V)^2 \\ & + 2\partial_x^2 [(\partial_x^3 V)(\partial_x V) \cdot] - 4\partial_x^2 [(\partial_x^2 V)^2 \cdot] + 4(\partial_x^3)(\partial_x V) \partial_x^2 - 8(\partial_x^2)^2 \partial_x^2 \} \\ & + \frac{1}{72} \tau^5 \varepsilon^{-1} (\partial_x^2 V)(\partial_x V)^2 - \frac{1}{630} \tau^7 \varepsilon^{-1} (\partial_x^3 V)(\partial_x V)^3 \\ & + \frac{1}{945} \tau^7 \varepsilon^{-1} \{ 2(\partial_x^3 V)(\partial_x V)^3 + 4(\partial_x^2 V)^2 (\partial_x V)^2 \} \\ & + \frac{101}{15120} \tau^7 \varepsilon^{-1} \{ (\partial_x^3 V)(\partial_x V)^3 + 2(\partial_x^2 V)^2 (\partial_x V)^2 \} \\ & + \frac{89}{3780} \tau^7 \varepsilon^{-1} (\partial_x^2 V)^2 (\partial_x V)^2 + \mathcal{O}(\varepsilon^{7/2}). \end{aligned}$$

6 Arranging in increasing powers of ε :

$$\begin{aligned}
 & \text{sBCH}(-R_0, Y_0) \\
 &= \overbrace{\tau \varepsilon \partial_x^2 + \frac{1}{12} \tau^2 \varepsilon^{-1} (\partial_x V)^2}^{\varepsilon^{1/2}} \\
 &+ \overbrace{\frac{1}{360} \tau^5 \varepsilon^{-1} (\partial_x^2 V) (\partial_x V)^2 - \frac{1}{6} \tau^3 \varepsilon \{ \partial_x^2 [(\partial_x^2 V) \cdot] + (\partial_x^2 V) \partial_x^2 \}}^{\varepsilon^{3/2}} \\
 &+ \overbrace{\frac{1}{12} \tau^3 \varepsilon (\partial_x^4 V)}^{\varepsilon^{5/2}} \\
 &+ \overbrace{\frac{1}{180} \tau^5 \varepsilon \{ -5 \partial_x^2 [(\partial_x^3 V) (\partial_x V) \cdot] + 4 \partial_x^2 [(\partial_x^2 V)^2 \cdot] - 5 (\partial_x^3) (\partial_x V) \partial_x^2 + 4 (\partial_x^2 V)^2 \partial_x^2 \}}^{\varepsilon^{5/2}} \\
 &+ \overbrace{\frac{1}{15120} \tau^7 \varepsilon^{-1} \{ 109 (\partial_x^3 V) (\partial_x V)^3 + 622 (\partial_x^2 V)^2 (\partial_x V)^2 \}}^{\varepsilon^{5/2}} + \mathcal{O}(\varepsilon^{7/2}).
 \end{aligned}$$

Note that the expansion now starts from $\mathcal{O}(\varepsilon^{1/2})$: we have got rid of the $\mathcal{O}(\varepsilon^{-1/2})$ term!

We now set

$$R_1 = \frac{1}{2}\tau\varepsilon\partial_x^2 + \frac{1}{24}\tau^3\varepsilon^{-1}(\partial_x V)^2, \quad Y_1 = \text{sBCH}(-R_0, Y_0).$$

We repeat steps 1–6 and this results in

$$\begin{aligned} \text{sBCH}(-R_1, Y_1) = & \overbrace{\frac{1}{360}\tau^5\varepsilon^{-1}(\partial_x^2 V)(\partial_x V)^2 - \frac{1}{6}\tau^3\varepsilon\{\partial_x^2[(\partial_x^2 V)\cdot] + (\partial_x^2 V)\partial_x^2\}}^{\varepsilon^{3/2}} \\ & + \overbrace{\frac{1}{12}\tau^3\varepsilon(\partial_x^4 V)}^{\varepsilon^{5/2}} \\ & + \overbrace{\frac{1}{180}\tau^5\varepsilon\{-\partial_x^2[(\partial_x^3 V)(\partial_x V)\cdot] + 8\partial_x^2[(\partial_x^2 V)^2\cdot] - (\partial_x^3)(\partial_x V)\partial_x^2 + 8(\partial_x^2 V)^2\partial_x^2\}}^{\varepsilon^{5/2}} \\ & + \overbrace{\frac{1}{7560}\tau^7\varepsilon^{-1}\{9(\partial_x^3 V)(\partial_x V)^3 + 8(\partial_x^2 V)^2(\partial_x V)^2\}}^{\varepsilon^{5/2}} + \mathcal{O}(\varepsilon^{7/2}). \end{aligned}$$

Thus, we have disposed of the $\mathcal{O}(\varepsilon^{1/2})$ terms: just one step, knocking out the $\mathcal{O}(\varepsilon^{3/2})$ terms, and we'll be done!

Thus, finally, R_2 is half the leading term in the expansion and

$$T_3 = Y_2 = \text{sBCH}(-R_1, Y_1).$$

To sum up,

$$R_0 = \frac{1}{2}\tau\varepsilon^{-1}V = \mathcal{O}(\varepsilon^{-1/2}),$$

$$R_1 = \frac{1}{2}\tau\varepsilon\partial_x^2 + \frac{1}{24}\tau^3\varepsilon^{-1}(\partial_x V)^2 = \mathcal{O}(\varepsilon^{1/2}),$$

$$R_2 = -\frac{1}{12}\tau^3\varepsilon\{\partial_x^2[(\partial_x^2 V)\cdot] + (\partial_x^2 V)\partial_x^2\} + \frac{1}{120}\tau^5\varepsilon^{-1}(\partial_x^2 V)(\partial_x V)^2 \\ = \mathcal{O}(\varepsilon^{3/2}),$$

$$T_3 = \frac{1}{12}\tau^3\varepsilon(\partial_x^4 V) + \frac{1}{180}\tau^5\varepsilon\{-\partial_x^2[(\partial_x^3 V)(\partial_x V)\cdot] \\ + 8\partial_x^2[(\partial_x^2 V)^2\cdot] - (\partial_x^3)(\partial_x V)\partial_x^2 + 8(\partial_x^2 V)^2\partial_x^2\} \\ + \frac{1}{7560}\tau^7\varepsilon^{-1}\{9(\partial_x^3 V)(\partial_x V)^3 + 8(\partial_x^2 V)^2(\partial_x V)^2\} = \mathcal{O}(\varepsilon^{5/2}).$$

We have our asymptotic splitting!

Once we discretise using **nodal values**, e^{R_0} is a diagonal matrix, while matrix/vector products with R_1, R_2 and T_3 can be computed by FFTs as part of a Krylov-subspace-based computation of their exponentials. If we discretise using **Fourier expansions**, also e^{R_0} is computed with FFT.

Stability

Suppose that each $(2r)$ th derivative is replaced by the **Hermitian** differentiation matrix \mathcal{K}_{2r} . Since $\tau = i\Delta t$ always features with an odd power, \mathcal{K}_{2r} , as well as the diagonal matrix that we obtain from discretizing multiplication operators, are multiplied by $\pm i$ and become **skew-Hermitian**. For example

$$R_0 \rightsquigarrow \frac{1}{2}(\Delta t)\varepsilon^{-1}i\mathcal{D}_V,$$

$$R_1 \rightsquigarrow \frac{1}{2}(\Delta t)\varepsilon i\mathcal{K}_2 - \frac{1}{24}(\Delta t)^3\varepsilon^{-1}i\mathcal{D}_{V_x^2}.$$

However... Because of replacement of odd derivatives, we have expressions of the form $i\mathcal{D}\mathcal{H}$ and $i\mathcal{H}\mathcal{D}$ where both \mathcal{D} and \mathcal{H} are **Hermitian** – and they are not **skew-Hermitian**! **Is this a problem?**

Not at all! Recalling that each term in \mathfrak{F} is multiplied by $\pm i$, we note that $\text{FLA}(i\partial_x^2, iV) \subset \mathfrak{su}(\mathbb{C})$. All our operations, using sBCH but also the replacement of odd derivatives, are Lie-algebra-compliant (linear combinations and commutators), therefore **everything** we have obtained in the course of our algorithm stays within $\mathfrak{su}(\mathbb{C})$.

As a sanity check, \mathcal{K} symmetric, \mathcal{D} diagonal $\Rightarrow i(\mathcal{K}\mathcal{D} + \mathcal{D}\mathcal{K})$ skew-Hermitian, hence

$$\begin{aligned}
 R_2 &\sim \frac{1}{12}(\Delta t)^3 \varepsilon i(\mathcal{K}_2 \mathcal{D}_{V_x^2} + \mathcal{D}_{V_x^2} \mathcal{K}_2) + \frac{1}{120}(\Delta t)^5 \varepsilon^{-1} i \mathcal{D}_{V_{xx} V_x^2} \\
 T_3 &\sim \frac{1}{12} \tau^3 \varepsilon \mathcal{D}_{\partial_x^4 V} + \frac{1}{180} \tau^5 \varepsilon \{8(\mathcal{K}_2 \mathcal{D}_{(\partial_x^2 V)^2} + \mathcal{D}_{(\partial_x^2 V)^2} \mathcal{K}_2) \\
 &\quad - (\mathcal{K}_2 \mathcal{D}_{(\partial_x^3 V)(\partial_x V)} + \mathcal{D}_{(\partial_x^3 V)(\partial_x V)} \mathcal{K}_2)\} \\
 &\quad + \frac{1}{7560} \tau^7 \varepsilon^{-1} \{9 \mathcal{D}_{(\partial_x^3 V)(\partial_x V)^3} + 8 \mathcal{D}_{(\partial_x^2 V)^2 (\partial_x V)^2}\}
 \end{aligned}$$

and it is easy to verify that both R_2 and T_3 are skew-Hermitian.

THEOREM It is true that $R_k \in \mathfrak{su}(\mathbb{C})$, $k = 0, 1, \dots, s$, and $T_{s+1} \in \mathfrak{su}(\mathbb{C})$.

LEADING WITH THE DERIVATIVE

Instead of $R_0 = \tau \varepsilon^{-1} V$, we start from $R_0 = \tau \varepsilon \partial_x^2$. On the face of it, this makes no sense, because we want to eliminate first the $\mathcal{O}(\varepsilon^{-1/2})$ term. However, let's try...

$$R_0 = \frac{1}{2} \tau \varepsilon \partial_x^2 = \mathcal{O}(\varepsilon^{1/2}),$$

$$R_1 = \frac{1}{2} \tau \varepsilon^{-1} V = \mathcal{O}(\varepsilon^{-1/2}),$$

$$R_2 = \frac{1}{12} \tau^3 \varepsilon^{-1} (\partial_x V)^2 = \mathcal{O}(\varepsilon^{1/2}),$$

$$R_3 = \frac{1}{24} \tau^3 \varepsilon \{ \partial_x^2 [(\partial_x^2 V) \cdot] + (\partial_x^2 V) \partial_x^2 \} + \frac{7}{240} \tau^5 \varepsilon^{-1} (\partial_x^2 V) (\partial_x V)^2 \\ = \mathcal{O}(\varepsilon^{3/2}),$$

$$T_4 = \frac{1}{24} \tau^3 \varepsilon (\partial_x^4 V) - \tau^5 \varepsilon \{ \frac{1}{30} \partial_x^2 [(\partial_x^3 V) (\partial_x V) \cdot] + \frac{1}{30} (\partial_x^3 V) (\partial_x V) \partial_x^2 \\ - \frac{1}{60} \partial_x^2 [(\partial_x^2 V)^2 \cdot] - \frac{1}{60} (\partial_x^2 V)^2 \partial_x^2 \} - \tau^7 \varepsilon^{-1} \{ \frac{163}{25704} (\partial_x^3 V) (\partial_x V)^3 \\ + \frac{211}{12852} (\partial_x^2 V)^2 (\partial_x V)^2 \} = \mathcal{O}(\varepsilon^{5/2}).$$

On the face of it, more terms. However... With nodal values e^{R_0} can be computed with a single FFT, R_1 and R_2 are diagonal (the other way around with Fourier expansions)—only $R_3 = \mathcal{O}(\varepsilon^{3/2})$ and $T_4 = \mathcal{O}(\varepsilon^{5/2})$ require Krylov subspace methods to compute their exponential!

SPACE DISCRETIZATION

In principle, we have three standard options to discretise ∂_x^{2s} to spectral accuracy:

- **Spectral methods:** We project in L_2 the solution on N th-degree trigonometric polynomials: the unknowns are *Fourier coefficients*, \mathcal{K} is diagonal and \mathcal{D} a circulant.
- **Spectral collocation:** We interpolate at equally-spaced points by N th-degree trigonometric polynomials: the unknowns are *nodal values*, \mathcal{K} is a circulant and \mathcal{D} is diagonal.
- **Pseudospectral method:** We wrap around a finite-difference method an infinite number of times: again, the unknowns are *nodal values*, \mathcal{K} is a circulant and \mathcal{D} is diagonal.

Except that, having committed already an $\mathcal{O}(\varepsilon^{7/2})$ error, we don't need spectral accuracy!

FINITE DIFFERENCES

We discretise in space to the same $\mathcal{O}(\varepsilon^{7/2}) = \mathcal{O}((\Delta x)^7)$ (actually, $\mathcal{O}((\Delta x)^8)$) accuracy using finite differences: $u_m \approx u(m\Delta x, \cdot)$,
 $m = -N, \dots, N$, $\Delta x = \frac{1}{N+\frac{1}{2}}$, and

$$u_m'' \approx \frac{1}{(\Delta x)^2} \left(-\frac{1}{560}u_{m-4} + \frac{8}{315}u_{m-3} - \frac{1}{5}u_{m-2} + \frac{8}{5}u_{m-1} - \frac{205}{72}u_m \right. \\ \left. + \frac{8}{5}u_{m+1} - \frac{1}{5}u_{m+2} + \frac{8}{315}u_{m+3} - \frac{1}{560}u_{m+4} \right).$$

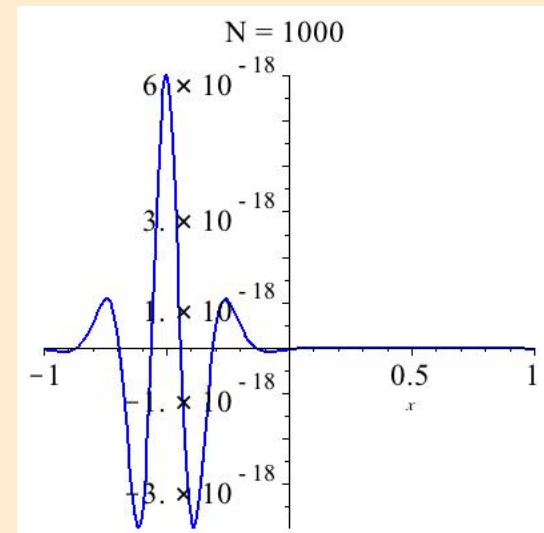
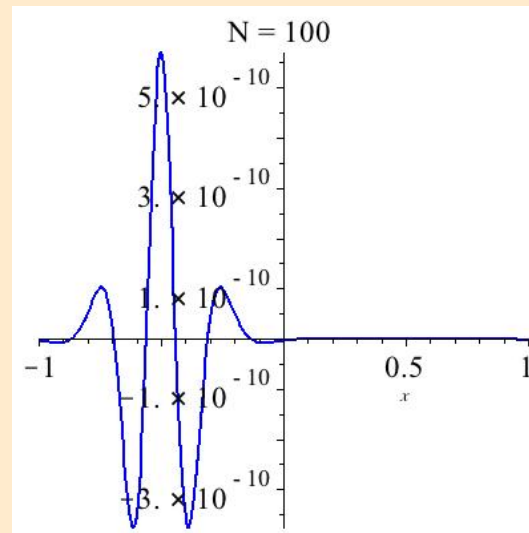
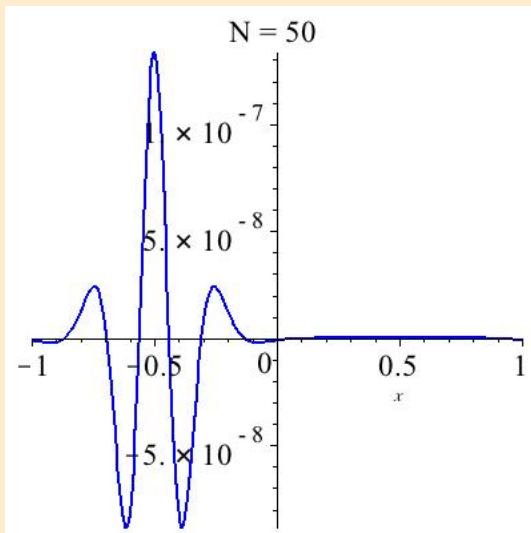
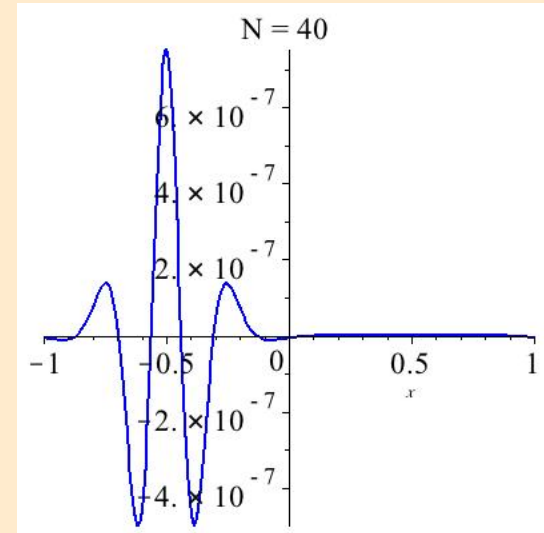
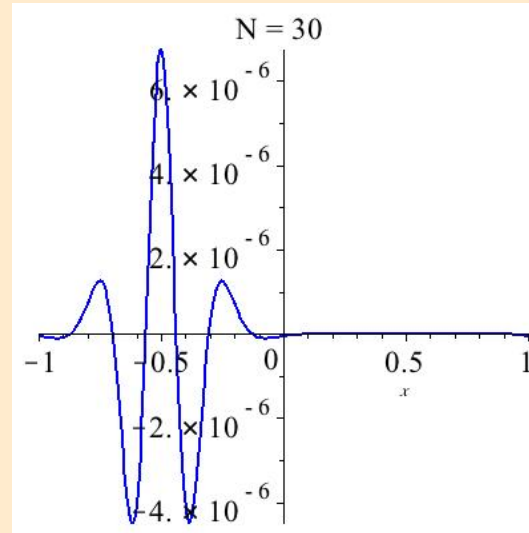
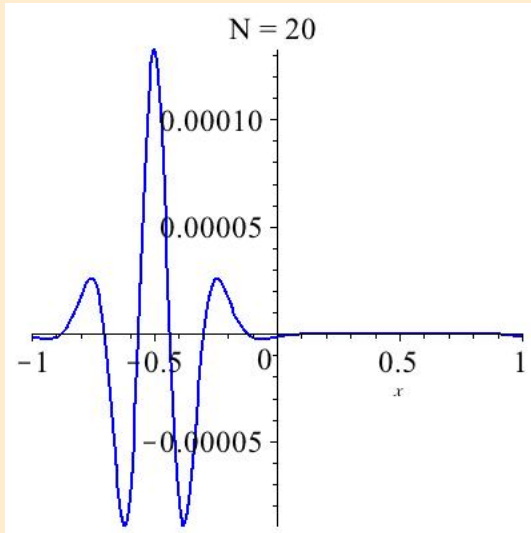
Therefore:

! \mathcal{K} is a 9-diagonal circulant, and

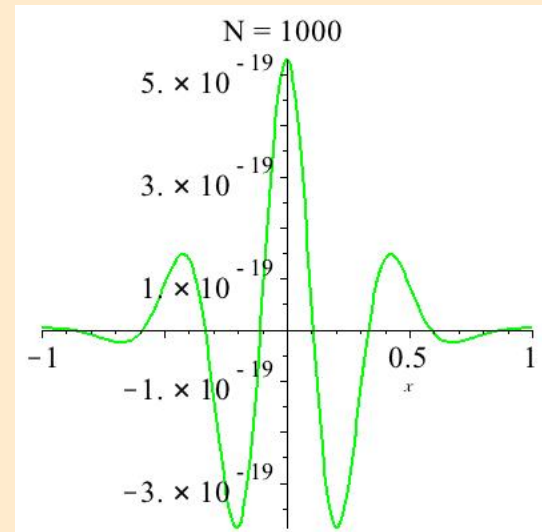
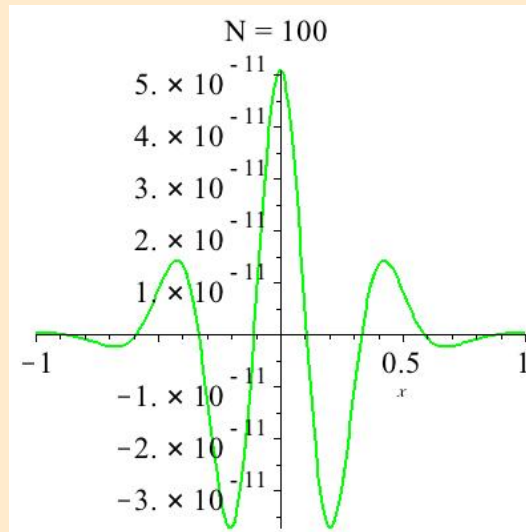
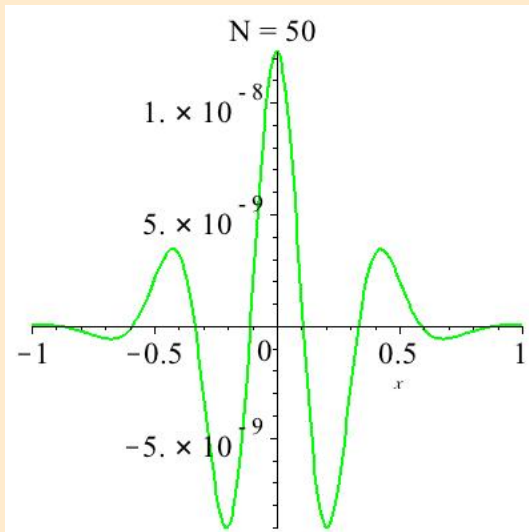
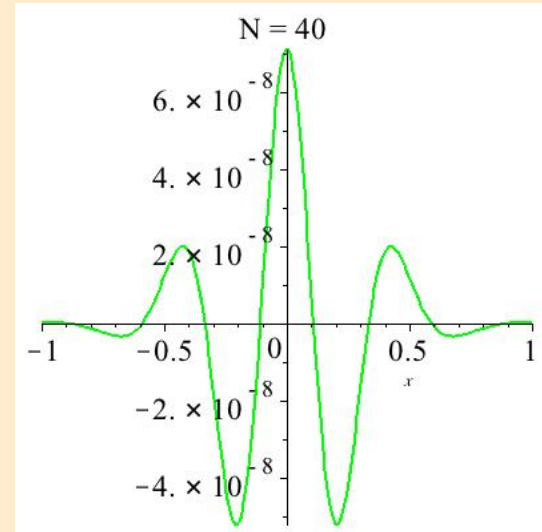
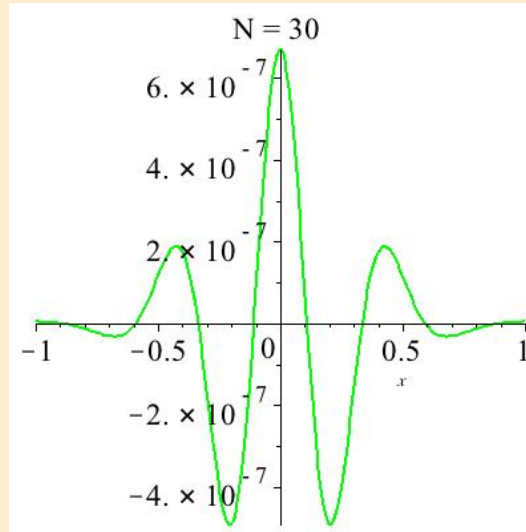
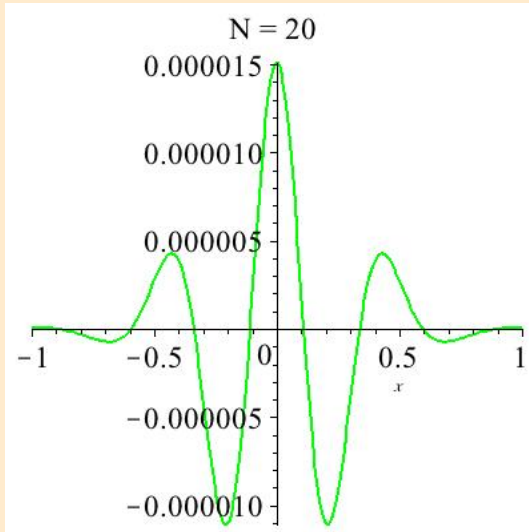
! \mathcal{D} is a diagonal matrix.

In that case all matrix-vector products are $\mathcal{O}(N)$ operations!

$$u(x) = \frac{1}{2 + \sin \pi x}$$



$$u(x) = e^{\cos \pi x}$$



So, does the error converge to a multiple of some mystery function?

Yes: in general, for the r th finite difference

$$u''(x) \approx \frac{1}{(\Delta x)^2} \sum_{m=-r}^r a_m u(x + m\Delta x)$$

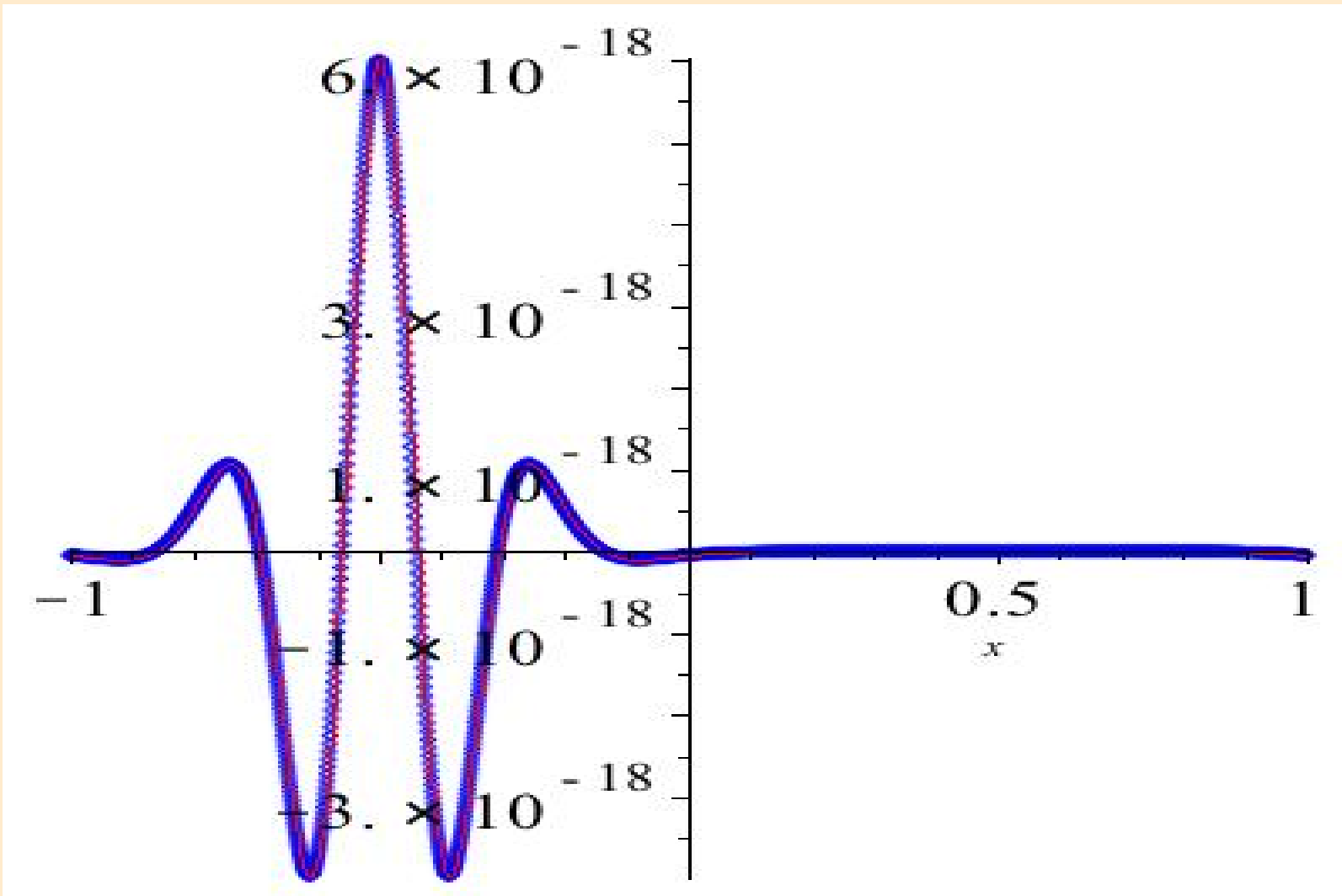
of order $2r$ we have

$$\sum_{m=-r}^r a_m \cos m\theta = -\theta^2 + c_r \theta^{2r+2} + \mathcal{O}(\theta^{2r+4}),$$

whereby the error is $c_r (\Delta x)^{2r} u^{(2r+2)}(x) + \mathcal{O}((\Delta x)^{2r+2})$. In our case

$c_4 = -\frac{1}{3150}$, therefore the error is

$$-\frac{1}{3150} (\Delta x)^8 u^{(10)}(x) + \mathcal{O}((\Delta x)^{10}), \quad \Delta x = \frac{1}{N + \frac{1}{2}}.$$



The error (in blue) and the error estimate (in red) for $u(x) = (2 + \sin \pi x)^{-1}$ and $N(= \mathcal{O}(\varepsilon^{1/2})) = 1000$.

COMPUTING THE EXPONENTIALS

Our method produces a dichotomy:

Arguments of the exponentials are either *trivial* or *small*!

They are *trivial* if the matrix is diagonal or is a Toeplitz circulant – in both cases computing the exponential is straightforward.

They are *small* if the argument is $A = \mathcal{O}(\varepsilon^\alpha)$ for some $\alpha > 0$. In that case the m th term in the Krylov basis

$$\{v, Av, A^2v, A^3v \dots\}$$

is $\mathcal{O}(\varepsilon^{m\alpha})$. Therefore, using estimates of Hochbruck, Lubich & Selhofer, it is possible to prove that we need *ridiculously* small dimension of a Krylov subspace to compute e^{Av} to $\mathcal{O}(\varepsilon^{7/2})$, say.

Eliminating the derivative first, $R_3 = \mathcal{O}(\varepsilon^{3/2})$ and $T_4 = \mathcal{O}(\varepsilon^{5/2})$ require dimensions 3 and 2 respectively!

CHALLENGES

- **Other pairs (ρ, σ) .** While the generalisation is straightforward (as a concept: the computation is fiendish), the interesting bit is to explore what are good choices of $\rho, \sigma \in (0, 1)$.
- **Time-dependent potentials.** Here we need to marry **Magnus expansions** with symmetric Zassenhaus. Preliminary results indicate that, although the algebra is formidable, this can be done: the underlying free Lie algebra is much more complicated but still embedded in \mathfrak{G} and “girth restrictions” apply.
- **Nonlinear Schrödinger.** Can we extend our method to the nonlinear Schrödinger equation $i\hbar u_t = -\frac{\hbar^2}{2m}\nabla^2 u - V(x)u + \lambda|u|^2u$? Straightforward generalisation does not work but is there a clever way this set of ideas can be extended?
- **Other equations.** Can this work with the wave equation? With its generalisations, e.g. Klein–Gordon? With Hamiltonian ODEs of the form $H(\mathbf{p}, \mathbf{q}) = \varepsilon H_1(\mathbf{p}, \mathbf{q}) + H_2(\mathbf{p}, \mathbf{q})$, ubiquitous in celestial mechanics?

